

# End of time in vector field theories

Fethi M Ramazanoğlu (with A Coates)  
Koç University

Star-UBB Seminar Series in Gravitation

December 8, 2022

# Summary

---

- Self interaction: Anything beyond the kinetic and mass terms. Example: Higgs.
- Main message: Self-interacting vector-fields are not physical in the strict sense
- Time evolution is possible, but not indefinitely: Problem is not detectable by “local” methods.
- Related to the “conditions” on the vectors: Likely goes beyond vectors.

Coates and Ramazanoglu, PRL 2022; [arXiv:2211.08027 \[gr-qc\]](https://arxiv.org/abs/2211.08027)

Closely related: Clough et al, PRL 2022; Mou and Zhang, PRL 2022

Inspiration: Silva et al PRD 2022; Demirboga et al PRD 2022; Garcia-Saenz et al PRL 2021

Earlier work: Esposito-Fareése et al PRD 2010

# Massless scalars

---

$$-\partial_t^2 \phi + \partial_x^2 \phi = 0 \quad (1)$$

$$\phi = e^{-i\omega t} e^{ikx} \Rightarrow \omega = \pm k \rightarrow \phi = e^{ikx} e^{\pm ikt}$$

Fourier modes oscillate in time.

# Massive scalars

---

$$-\partial_t^2 \phi + \partial_x^2 \phi = m^2 \phi \quad (2)$$

$$\phi = e^{-i\omega t} e^{ikx} \Rightarrow \omega = \pm \sqrt{k^2 + m^2} \rightarrow \phi = e^{ikx} e^{\pm i\sqrt{k^2 + m^2} t}$$

Fourier modes *still* oscillate in time.

# Massive scalars: alternative look

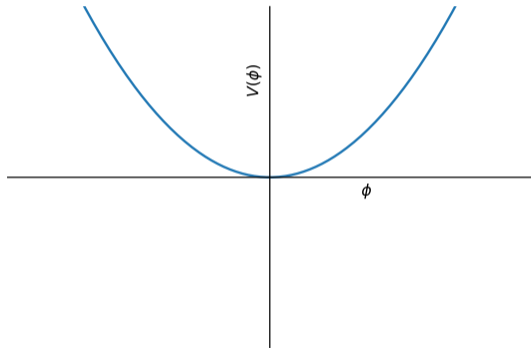
---

$$-\partial_t^2 \phi + \partial_x^2 \phi = V'(\phi), \quad V(\phi) = \frac{1}{2} m \phi^2 \quad (3)$$

$$\partial_x \phi = 0 \Rightarrow \ddot{\phi} = -V'(\phi) \quad (4)$$

$$\Rightarrow \frac{1}{2} \dot{\phi}^2 + \frac{1}{2} m^2 \phi^2 = \text{const} \quad (5)$$

Particle moving under the potential  $V(\phi)$ .



# Self-interacting scalars

---

$$-\partial_t^2 \phi + \partial_x^2 \phi = V'(\phi) \quad (6)$$

$$V(\phi) = \frac{1}{2} m^2 \phi^2 + \frac{1}{4} \lambda m^2 \phi^4 \quad (7)$$

$\lambda > 0 \rightarrow$  fine

$\lambda < 0 \rightarrow$  potential trouble

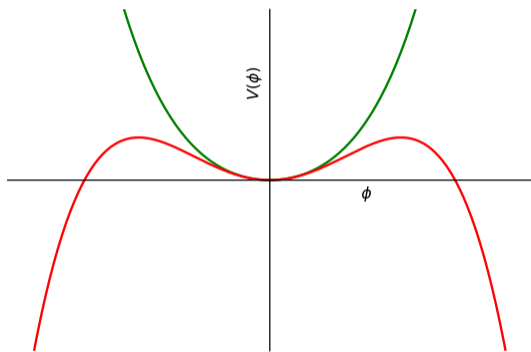


Figure:  $V(\phi) = \frac{1}{2} \phi^2 \pm \frac{1}{4} \phi^4$

## Changing the derivative terms – Ill-posedness

---

$$+\partial_t^2\phi + \partial_x^2\phi = 0 \quad (8)$$

$$\phi = e^{-i\omega t} e^{ikx} \Rightarrow \omega = \pm ik \rightarrow \phi = e^{ikx} e^{|k|t}$$

Arbitrarily fast blow up for  $k \rightarrow \infty!$

There are modes that *immediately* go to infinity if we try time evolution.

$$\begin{aligned} \phi(0, x=0), \quad \partial_t\phi(0, x) &= \epsilon e^{i\frac{x}{\epsilon}} \\ \Rightarrow \phi(x, t) &= \epsilon^2 e^{i\frac{x}{\epsilon}} \sinh \frac{t}{\epsilon} \rightarrow \text{no continuous dependence!} \end{aligned} \quad (9)$$

# General case

---

$$A \partial_t^2 \phi + 2B \partial_t \partial_x \phi + C \partial_x^2 \phi = g^{\mu\nu} \nabla_\mu \nabla_\nu \phi \Rightarrow g^{\mu\nu} = \begin{bmatrix} A & B \\ B & C \end{bmatrix} \quad (10)$$

Well-posed if eigenvalues are  $(-, +, \dots, +)$ : **Signature of the metric.**

Lower derivative terms are not critical.



# Scalar field summary

---

$$g^{\mu\nu}\nabla_{\mu}\nabla_{\nu}\phi = V'(\phi) \tag{11}$$

- $V(\phi)$  unbounded from below: Bad physics, math *maybe* OK.
- Metric signature not Lorentzian: ill posed, even math is in trouble.
- The highest derivative terms determine well-posedness, the metric has to be Lorentzian. Good thing is that the metric is a given ... or is it?

# Massless vectors

---

$$S = - \int d^n x \sqrt{|g|} F_{\mu\nu} F^{\mu\nu} \quad (12)$$

$$F_{\mu\nu} = \nabla_\mu X_\nu - \nabla_\nu X_\mu = \partial_\mu X_\nu - \partial_\nu X_\mu \quad (13)$$

$$0 = \overbrace{\eta^{\mu\nu} \nabla_\mu \nabla_\nu}^{-\partial_t^2 + \partial_x^2} X^\rho - \nabla^\rho \nabla_\mu X^\mu = \nabla_\mu F^{\mu\rho} \quad (14)$$

$X^\mu$  has gauge freedom! We can *choose*  $X^\mu$  such that  $\nabla_\mu X^\mu = \partial_t \Phi + \nabla \cdot \vec{A} = 0 \rightarrow$  Lorenz gauge.

$$0 = \eta^{\mu\nu} \nabla_\mu \nabla_\nu X^\rho \quad (15)$$

# Massive vectors

---

$$S = - \int d^n x \sqrt{|g|} [F_{\mu\nu} F^{\mu\nu} + 2m^2 X_\mu X^\mu] \quad (16)$$

$$F_{\mu\nu} = \nabla_\mu X_\nu - \nabla_\nu X_\mu = \partial_\mu X_\nu - \partial_\nu X_\mu \quad (17)$$

$$m^2 X^\rho = \eta^{\mu\nu} \nabla_\mu \nabla_\nu X^\rho - \nabla^\nu \nabla_\mu X^\mu \quad (18)$$

$$\nabla_\rho \nabla_\mu F^{\mu\rho} = m^2 \nabla_\rho X^\rho \quad (19)$$

$$\Rightarrow \nabla_\mu X^\mu = 0 \rightarrow \text{Lorenz condition, not gauge!} \quad (20)$$

# Self-interacting vectors

---

$$S = - \int d^n x \sqrt{|g|} [F_{\mu\nu} F^{\mu\nu} + 2m^2 X_\mu X^\mu + \lambda m^2 (X_\mu X^\mu)^2] \quad (21)$$

$$m^2(1 + \lambda X^2)X^\rho = \eta^{\mu\nu} \nabla_\mu \nabla_\nu X^\rho - \nabla^\rho \nabla_\mu X^\mu \quad (22)$$

$$\nabla_\rho \nabla_\mu F^{\mu\rho} = m^2 \nabla_\rho [(1 + \lambda X^2)X^\rho] \quad (23)$$

$$\Rightarrow \nabla_\rho [(1 + \lambda X^2)X^\rho] = 0 \rightarrow \text{No easy cancellation} \quad (24)$$

# Self-interacting vectors: Principal part

---

$$m^2(1 + \lambda X^2)X^\rho = \eta^{\mu\nu} \nabla_\mu \nabla_\nu X^\rho - \nabla^\rho \nabla_\mu X^\mu \quad (25)$$

... generalized Lorenz condition+algebra ...

$$\Rightarrow 0 = [(1 + \lambda X^2)\eta^{\mu\nu} + 2\lambda X^\mu X^\nu] \nabla_\mu \nabla_\nu X^\rho + \dots \quad (26)$$

$$\Rightarrow 0 = \bar{g}^{\mu\nu} \partial^\mu \partial^\nu X^\alpha + \dots \quad (27)$$

$$\bar{g}^{\mu\nu} = (1 + \lambda X^2) \begin{bmatrix} -1 + \frac{2\lambda X^t X^t}{1 + \lambda X^2} & \frac{2\lambda X^t X^x}{1 + \lambda X^2} \\ \frac{2\lambda X^t X^x}{1 + \lambda X^2} & 1 + \frac{2\lambda X^x X^x}{1 + \lambda X^2} \end{bmatrix} \text{ governs the dynamics, not } \eta^{\mu\nu}!$$

Even if the spacetime is perfectly OK,  $\bar{g}_{\mu\nu}(X^\mu)$  can have the wrong signature!

# Self-interacting vectors: Principal part

---

- Works the same way for a general curved spacetime  $\eta_{\mu\nu} \rightarrow g_{\mu\nu}$ .
- However, curvature plays no essential role, *intrinsic* to the vector field theory, i.e. even occurs for fixed  $g_{\mu\nu} = \eta_{\mu\nu}$ .
- More technical machinery is needed for higher than  $1 + 1D$ , but the same result.

## Problematic $\bar{g}_{\mu\nu}$

---

Singular metric when determinant vanishes

$$\bar{g} = \det \bar{g}_{\mu\nu} = g (1 + \lambda X^2)^d (1 + 3\lambda X^2) = g z^d z_3 = 0 \quad (28)$$

$$\Rightarrow X^2 = -\lambda/3 \quad (29)$$

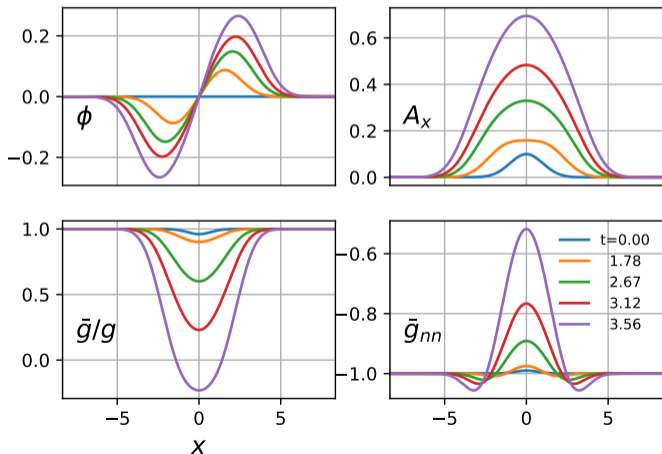
$\bar{g}_{\mu\nu}$  becomes singular at finite values of  $X^\mu$

confirmed by checking the divergence of curvature

# Dynamical loss of hyperbolicity

Start with Lorentzian  $\bar{g}_{\mu\nu}$ , does the problem occur?

Yes, for any value of  $m, \lambda$ .

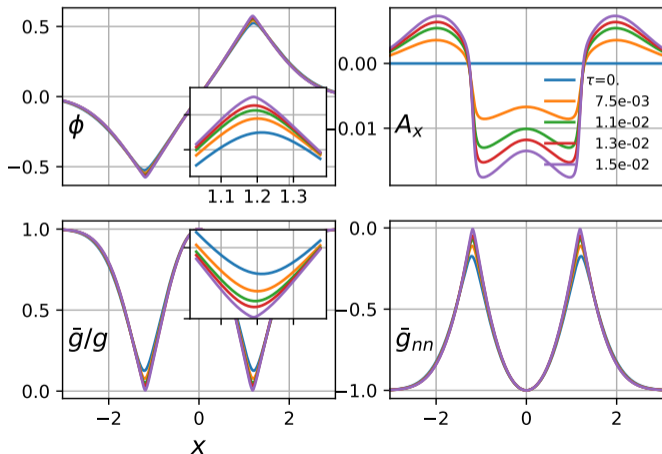




# Dynamical loss of hyperbolicity

Start with Lorentzian  $\bar{g}_{\mu\nu}$ , does the problem occur?

Yes, for any value of  $m, \lambda$ .



# On progress, future

---

- Likely holds for anything beyond Proca.
- Likely holds for other theories:  $p$ -form fields. **New theoretical test!**
- Backreaction effects not known.
- Similar problems in nonminimal couplings, tachyon on vector  $\Rightarrow$  ill-posedness.
- Might be overlooked/resolved when the theory is *effective*.

# Effective field theories, fixing-the-equation

---

Abelian Higgs theory: define  $D_\mu\phi = \partial_\mu\phi - iqA_\mu\phi$ .

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2} \left[ (\overline{D_\mu\phi}) (D^\mu\phi) - m^2|\phi|^2 + \frac{\Lambda}{2}|\phi|^4 \right] \quad (30)$$

$$\Rightarrow \nabla_\mu F^{\mu\nu} = -q\rho^2\partial^\nu\Theta + q^2\rho^2A^\nu \quad , \quad (\phi = \rho \exp(i\Theta)) \quad (31)$$

$$\Rightarrow \nabla_\mu F^{\mu\nu} = \frac{q^2 m^2}{\Lambda} \left( 1 - \frac{q^2 X^2}{m^2} \right) X^\nu + \mathcal{O} \left( \frac{1}{m^2} \right) \quad (32)$$

The first equation is fine, no loss of hyperbolicity.

Only known for  $\lambda < 0$ , other UV completion options?

# The end

---

QUESTIONS?

## Some wrong turns

---

Evolve the vector with 3 + 1 decomposition.

$$ds^2 = -\alpha^2 dt^2 + \gamma_{ij}(dx^i + \beta^i dt)(dx^j + \beta^j dt) \quad (33)$$
$$X_\mu = n_\mu \phi + A_\mu, \quad \phi = -n_\mu X^\mu, \quad A_i = (\delta^\mu_i + n^\mu n_i) X_\mu.$$

$$\begin{aligned} \partial_t \phi &= \beta^i D_i \phi - A^i D_i \alpha - \frac{\alpha}{\bar{g}_{nn}} z (K \phi - D_i A^i) \\ &+ \frac{2\lambda\alpha}{\bar{g}_{nn}} [A^i A^j D_i A_j - \phi (E_i A^i - K_{ij} A^i A^j + 2A^i D_i \phi)] \quad (34) \\ 0 &= D_i E^i + \mu^2 z \phi = \mathcal{C}, \end{aligned}$$

$$\bar{g}_{nn} = n^\mu n^\nu \bar{g}_{\mu\nu} = -(1 + 3\lambda X^2) + 2\lambda A_i A^i. \quad (35)$$

# A numerical look at ill-posedness

---

$$+\partial_t^2\phi + \partial_x^2\phi = 0 \quad (36)$$

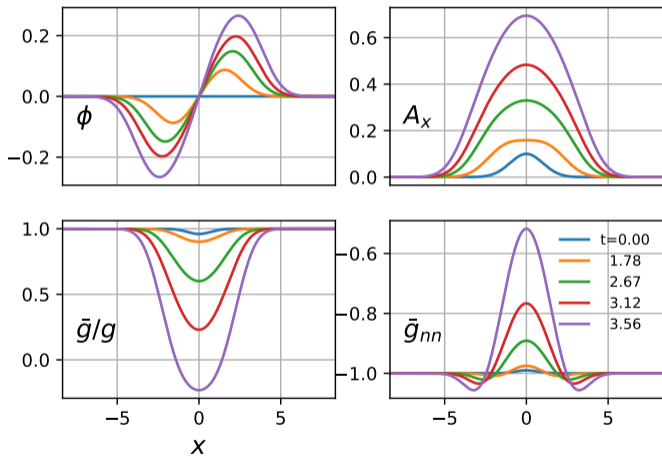
$$\phi = e^{-i\omega t} e^{ikx} \Rightarrow \omega = \pm ik \rightarrow \phi = e^{ikx} e^{|k|t}$$

Arbitrarily fast blow up for  $k \rightarrow \infty!$

There are modes that *immediately* go to infinity if we try time evolution.

$$\begin{aligned} \phi(0, x=0), \quad \partial_t\phi(0, x) &= \epsilon e^{i\frac{x}{\epsilon}} \\ \Rightarrow \phi(x, t) &= \epsilon^2 e^{i\frac{x}{\epsilon}} \sinh \frac{t}{\epsilon} \rightarrow \text{no continuous dependence!} \end{aligned} \quad (37)$$

# A numerical look at ill-posedness



# A numerical look at ill-posedness

