

# *On the Nonrelativistic Expansion of GR*

Jelle Hartong

University of Edinburgh, School of Mathematics

STAR-UBB Institute

Babeş-Bolyai University

30 March, 2023

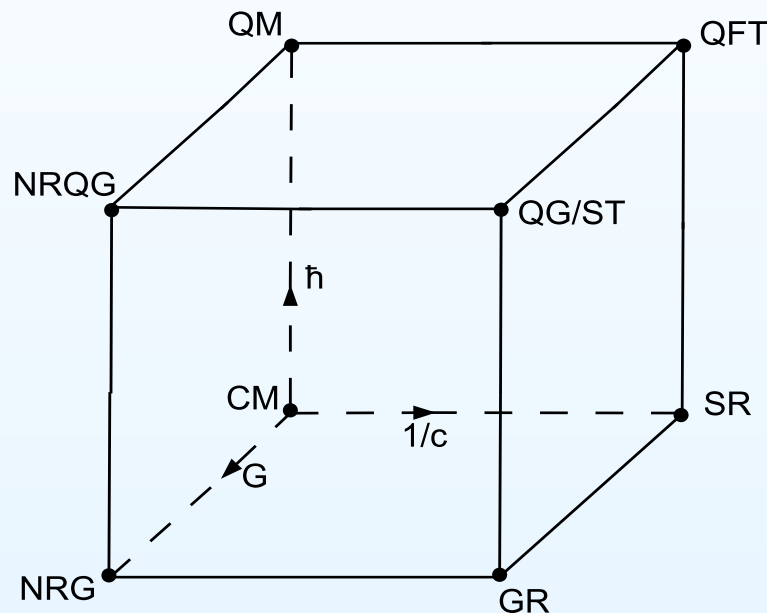
In various collaborations with:

Jørgen Musaeus (UoE)

Emil Have (UoE), Niels Obers (Nordita & NBI), Igor Pikovski (Stockholm University)

# Introduction

- Bronstein cube

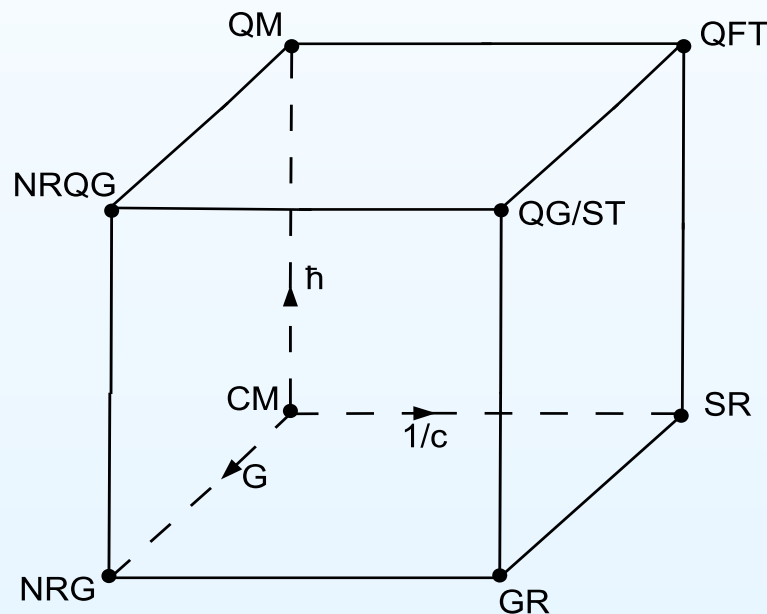


- **NRG = non-relativistic gravity**  $\supset$  Newtonian gravity
- $1/c$  expansion  $\supset$  post-Newtonian expansion
- **NRQG = non-relativistic quantum gravity**

- NRQG is a bit of a misnomer. There is no dynamical gravity to quantise. Think of quantum matter (described by quantum mechanics) backreacting with a background that reacts instantaneously.

# Introduction

- Is there a well-defined non-relativistic limit of quantum gravity/string theory?



- In string theory there exist dualities between QG and specific QFTs (holography, AdS/CFT, ...).
- Do such dualities exist in the NR domain, i.e. can NR strings be dual to some NR field theory?

- Limits of AdS/CFT called Spin Matrix Theory give rise to NR strings dual to quantum mechanical limits of AdS/CFT. See e.g. [Harmark, Kristjansson, Orselli, 2007/8], [Harmark, JH, Obers, 2017].

## Introduction: NR geometry

---

- Many of the recent developments in NR gravity, and NR string theory rely on an improved understanding of NR geometry.
- The most common example of a NR geometry is Newton–Cartan geometry which is the arena of every day life.
- Many other NR geometries have been found: type II Newton–Cartan, string Newton–Cartan, Aristotelian, Carrollian geometries ...
- Outside of GR and string theory, NR geometry has found applications in fluid dynamics, condensed matter physics, Hořava–Lifshitz gravity, 2D/3D gravity, ...

# Introduction: NR approximations of GR

---

- ‘Drawbacks’ of Post-Newtonian approximation methods (Blanchet–Damour and Will–Wiseman approaches):
  - harmonic gauge
  - strong no-incoming radiation boundary condition
  - compactly supported matter
- An approach to address the first two issues is currently WIP [JH, Musaeus].
- We do not have a universal method to define non-relativistic approximations of GR coupled to any matter system (compact or non-compact) in any gauge.
- The PN approximation is a weak field approximation. There is however also strong non-relativistic gravity. Can this be useful?

## Introduction: PN corrections to quantum mechanics

---

- A very special case of NRQG is to take  $G \rightarrow 0$  and study QM on a fixed background.
- Is there a coupling prescription for this?
- Suppose we know the  $1/c$  corrections from SR how do we couple the system to geometries that are obtained from  $1/c$  expansions of solutions of GR?
- What is the dynamics of a hydrogen atom in a Kerr background to some order in  $1/c$ ?

# Outline

---

- Newton–Cartan geometry
- $1/c$  expansion of GR
- QM on NR geometries

# Newton–Cartan Geometry

$$\text{metric : } \tau(a, b) = |t' - t| = \int_a^b \tau, \quad \rho(a, b) = \|\vec{y} - \vec{x}\| = \int_a^b ds$$

- Here  $\tau = dt$  and  $ds^2|_{t=\text{cst}} = \delta_{ij}dx^i dx^j$
- Write  $\tau = \tau_\mu dx^\mu$  with coordinates  $x^\mu = (t, x^i)$ . Remove the restriction to  $t = \text{cst}$  in  $ds^2$  and write  $ds^2 = h_{\mu\nu}dx^\mu dx^\nu$  as a quadratic form with signature  $(0, 1, \dots, 1)$ .
- Under a Galilean boost with parameter  $\lambda_\mu = (0, \vec{v})$

$$h_{\mu\nu} \rightarrow h_{\mu\nu} + \lambda_\mu \tau_\nu + \lambda_\nu \tau_\mu + \lambda^2 \tau_\mu \tau_\nu$$

- A manifold's tangent space is the flat version of the manifold. In general  $\tau_\mu$ ,  $h_{\mu\nu}$  and the parameter  $\lambda_\mu$  are tensor fields.

$$\tau = dt \quad \text{absolute time}$$

$$\tau = N dt \quad \text{absolute foliation}$$



# Newton–Cartan Geometry

- Mass is like electric charge  $\Rightarrow$  gauge connection  $m_\mu$
- Fields on curved NC geometry

$$\delta S[\tau_\mu, h_{\mu\nu}, m_\mu] = \int d^4x e \left[ \mathcal{E}^\mu \delta\tau_\mu + \frac{1}{2} \mathcal{T}^{\mu\nu} \delta h_{\mu\nu} + J^\mu \delta m_\mu \right]$$

$\mathcal{E}^\mu$	energy current
$\mathcal{T}^{\mu\nu}$	momentum-stress tensor
$J^\mu$	mass current

- momentum = mass flux:  $\mathcal{T}^{0i} = J^i \Leftrightarrow \delta m_\mu = \lambda_\mu$  and  $\delta h_{\mu\nu} = \lambda_\mu \tau_\nu + \lambda_\nu \tau_\mu$
- mass conservation:  $\partial_\mu (e J^\mu) = 0 \Leftrightarrow \delta m_\mu = \partial_\mu \sigma$
- Triplet  $(\tau_\mu, h_{\mu\nu}, m_\mu)$  with  $\lambda_\mu, \sigma$  gauge redundancy defines a NC geometry.

# Newton–Cartan Geometry

- Geodesic in NC geometry: Newton's equation

$$S = m \int d\lambda \left( \frac{h_{\mu\nu} \dot{x}^\mu \dot{x}^\nu}{2\tau_\rho \dot{x}^\rho} - m_\mu \dot{x}^\mu \right)$$

- The time component of  $m_\mu$  is Newton's potential.
- The fact that the mass is only an overall coupling is a manifestation of the equivalence principle.
- Schrödinger wavefunction on NC geometry

$$S = \int d^4x e (im\psi^* v^\mu D_\mu \psi - im\psi v^\mu D_\mu \psi^* - h^{\mu\nu} D_\mu \psi D_\nu \psi^*)$$

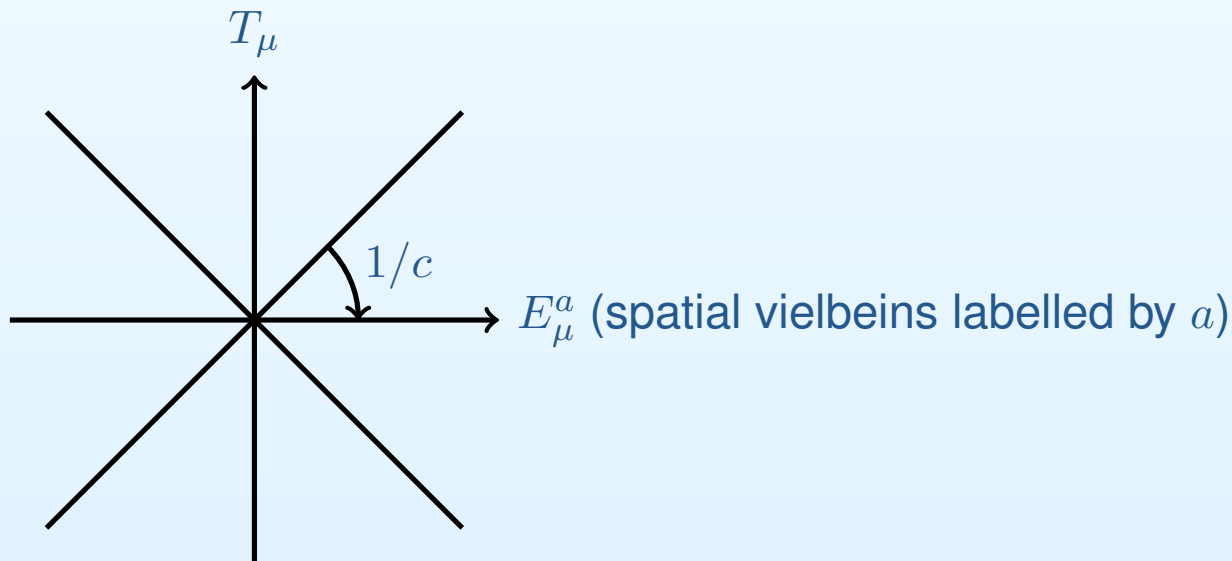
$v^\mu$ ,  $h^{\mu\nu}$  inverses of  $\tau_\mu$  and  $h_{\mu\nu}$ .

- $U(1)$  symmetry is gauged by  $m_\mu$ :  $D_\mu \psi = \partial_\mu \psi + im m_\mu \psi$ .
- Mass conservation = conservation of probability.

# 1/c expansion of GR

Review article: [JH, Obers, Oling, 2022]

- A convenient way to make the  $c$ -dependence of GR manifest is to write  $g_{\mu\nu} = -c^2 T_\mu T_\nu + \Pi_{\mu\nu}$  and  $g^{\mu\nu} = -\frac{1}{c^2} T^\mu T^\nu + \Pi^{\mu\nu}$ .
- Signature of  $\Pi_{\mu\nu}$  is  $(0, 1, \dots, 1)$ .
- Light cones in tangent space have slope  $1/c$ :



# 1/c expansion of GR

- So far we just reformulated GR in different variables. We will now assume that we can Taylor expand  $T_\mu$  and  $\Pi_{\mu\nu}$  in  $1/c$ :

$$T_\mu = \tau_\mu + \frac{1}{c^2} m_\mu + \frac{1}{c^4} B_\mu + \mathcal{O}(c^{-6}), \quad \Pi_{\mu\nu} = h_{\mu\nu} + \frac{1}{c^2} \Phi_{\mu\nu} + \mathcal{O}(c^{-4})$$

- This is what leads to the covariant  $1/c$  expansion.
- Note here only even powers. For odd powers see [Ergen, Hamamci, Van den Bleeken, 2020] and later in the PN expansion.
- This leads to the metric expansion:

$$g_{\mu\nu} = -c^2 \tau_\mu \tau_\nu + h_{\mu\nu} - 2\tau_{(\mu} m_{\nu)} + c^{-2} (\Phi_{\mu\nu} - m_\mu m_\nu - 2\tau_{(\mu} B_{\nu)}) + \mathcal{O}(c^{-4})$$

- The  $1/c$  expansion of the metric was pioneered by [Dautcourt, 1990/97] and generalised in [Van den Bleeken, 2017], [JH, Hansen, Obers, 2018-20].

# 1/c expansion of GR

- We can view the  $1/c$  expansion as an expansion around a geometry described by  $\tau_\mu$  and  $h_{\mu\nu}$  where all the higher order fields  $m_\mu$  and  $\Phi_{\mu\nu}$  are like gauge connections.
- Expanding the generator of infinitesimal diffeos:  
 $\Xi^\mu = \xi^\mu + \frac{1}{c^2}\zeta^\mu + O(c^{-4})$  leads to gauge transformations for the subleading fields  $m_\mu$  and  $\Phi_{\mu\nu}$  w.r.t. subleading diffeos  $\zeta^\mu$ .
- Local Lorentz transformations acting on  $T_\mu$  and  $\Pi_{\mu\nu}$  also get expanded and lead to local Galilean transformations.
- Expanding the Einstein equations coupled to a point particle leads to Newtonian gravity:

$$\bar{R}_{\mu\nu} = 8\pi G \frac{d-2}{d-1} \rho \tau_\mu \tau_\nu, \quad d\tau = 0$$

where we used the leading order Levi-Civita connection  $\bar{\Gamma}_{\mu\nu}^\rho$ .

# 1/c expansion of GR: Examples of weak limits

$$ds_{\text{Schwarzschild}}^2 = -c^2 \left(1 - \frac{2Gm}{c^2 r}\right) dt^2 + \left(1 - \frac{2Gm}{c^2 r}\right)^{-1} dr^2 + r^2 d\Omega_{S^2}^2$$

$$ds_{\text{AdS(+)/dS(-)}}^2 = -c^2 \left(1 \pm \frac{r^2}{l^2}\right) dt^2 + \frac{dr^2}{1 \pm \frac{r^2}{l^2}} + r^2 d\Omega_{S^2}^2$$

- Consider  $m$  independent of  $c$ .

$$\tau_\mu dx^\mu = dt, \quad h_{\mu\nu} dx^\mu dx^\nu = dr^2 + r^2 d\Omega_{S^2}^2, \quad m_\mu dx^\mu = -\frac{Gm}{r} dt$$

Point mass in flat spacetime with Newtonian pot.  $\Phi = -\frac{Gm}{r}$ .

- Take  $l = c/H$  with the Hubble constant  $H$  independent of  $c$ .

$$\tau_\mu dx^\mu = dt, \quad h_{\mu\nu} dx^\mu dx^\nu = dr^2 + r^2 d\Omega_{S^2}^2, \quad m_\mu dx^\mu = \pm \frac{1}{2} H^2 r^2 dt$$

# $1/c$ expansion of GR: Example of a strong limit

- Strong limit:  $m = c^2 M$ ;  $M$  independent of  $c^2$  [Van den Bleeken, 2017].

$$\tau_\mu dx^\mu = \sqrt{1 - \frac{2GM}{r}} dt, \quad h_{\mu\nu} dx^\mu dx^\nu = \left(1 - \frac{2GM}{r}\right)^{-1} dr^2 + r^2 d\Omega_{S^2}$$

$$m_\mu dx^\mu = 0 = \Phi_{\mu\nu} dx^\mu dx^\nu$$

- This strong gravity expansion of the Schwarzschild metric is not captured by Newtonian gravity, but is still described as a Newton–Cartan geometry.
- This provides us with a different approximation of GR as compared to the post-Newtonian expansion.
- $\tau$  is no longer exact but  $\tau \wedge d\tau = 0$  (hypersurface orthogonality). Strong limit captures gravitational time dilation: clocks tick slower/faster depending on position on a constant time slice.

# QM on NR geometries

---

[JH, Have, Obers, Pikovski, to appear]

- Goal: given the Schrödinger equation for a quantum system in flat space time, including  $1/c^2$  corrections, find a coupling prescription to couple this to  $1/c^2$  expanded geometries.
- A top-down approach for the Schwarzschild metric was used in [Lämmerzahl, 1995] and generalised in [Schwartz, Giulini, 2018].
- Here the focus is on coupling prescriptions (i.e. general backgrounds and systematising results).
- We ignore backreaction: classically this leads to the Schrödinger–Newton equation which violates the superposition principle.



# QM on NR geometries: The main idea

- KG equation and inner product on solution space:

$$-\frac{1}{c^2}\partial_t^2\phi + \nabla^2\phi - m^2c^2\phi = 0$$

$$\langle\phi_2|\phi_1\rangle = -\frac{i}{c^2}\int_{t=\text{cst}}d^3x(\phi_1\partial_t\phi_2^* - \phi_2^*\partial_t\phi_1)$$

- Define  $\phi = \frac{1}{\sqrt{2m}}e^{-imc^2t}\psi$  with  $\psi = \psi_{\text{LO}} + c^{-2}\psi_{\text{NLO}} + \dots$

$$-\frac{1}{c^2}\partial_t^2\psi + 2im\partial_t\psi + \nabla^2\psi = 0$$

$$\langle\phi_2|\phi_1\rangle = \int_{t=\text{cst}}d^3x\left[\psi_2^*\psi_1 - \frac{i}{2mc^2}(\psi_1\partial_t\psi_2^* - \psi_2^*\partial_t\psi_1)\right]$$

- Define  $\Psi = \psi - \frac{1}{4m^2c^2}\nabla^2\psi + \dots$  s.t.  $\langle\phi_2|\phi_1\rangle = \int_{t=\text{cst}}d^3x\Psi_2^*\Psi_1$

$$i\partial_t\Psi = -\frac{1}{2m}\nabla^2\Psi - \frac{1}{8m^3c^2}\nabla^4\Psi + O(c^{-4})$$

# QM on NR geometries: coupling prescription

- We assume for simplicity flat NC spacetime:  $\tau = dt$ ,  $h = dx^i dx^i$ .
- Again  $\phi = \frac{1}{\sqrt{2m}} e^{-imc^2 t} \psi$  where  $\psi = \psi_{\text{LO}} + c^{-2} \psi_{\text{NLO}} + \dots$
- Using  $\delta\phi = \Xi^\mu \partial_\mu \phi$  with  $\Xi^\mu = \xi^\mu + c^{-2} \zeta^\mu + c^{-4} \chi^\mu + \dots$  we find

$$\delta\psi_{\text{LO}} = -im\zeta^t \psi_{\text{LO}}$$

$$\delta\psi_{\text{NLO}} = -im\zeta^t \psi_{\text{NLO}} + \zeta^t \partial_t \psi_{\text{LO}} - im\chi^t \psi_{\text{LO}} + \zeta^i \partial_i \psi_{\text{LO}}$$

- The  $\zeta^t, \chi^t$  are gauge transformation parameters with gauge fields  $m_\mu, B_\mu$ , that appear in the expansion of the vielbeins.
- We will deal with  $\zeta^i$  separately.

## QM on NR geometries: coupling prescription

- The coupling prescription is a statement about how to couple

$$\text{EOM}_{\text{LO}} = 2im\partial_t\psi_{\text{LO}} + \nabla^2\psi_{\text{LO}} = 0$$

$$\text{EOM}_{\text{NLO}} = 2im\partial_t\psi_{\text{NLO}} + \nabla^2\psi_{\text{NLO}} - \partial_t^2\psi_{\text{LO}} = 0$$

to the gauge fields  $m_\mu$ ,  $B_\mu$ .

- The guiding principle is to find covariant derivatives such that

$$\delta\mathcal{D}_\mu\psi_{\text{LO}} = -im\zeta^t\mathcal{D}_\mu\psi_{\text{LO}}$$

$$\delta\mathcal{D}_\mu\psi_{\text{NLO}} = \zeta^t\partial_t\mathcal{D}_\mu\psi_{\text{LO}} - im\zeta^t\mathcal{D}_\mu\psi_{\text{NLO}} - im\chi^t\mathcal{D}_\mu\psi_{\text{LO}}$$

- This leads to

$$\mathcal{D}_\mu\psi_{\text{LO}} = \partial_\mu\psi_{\text{LO}} + imm_\mu\psi_{\text{LO}}$$

$$\mathcal{D}_\mu\psi_{\text{NLO}} = \partial_\mu\psi_{\text{NLO}} + imm_\mu\psi_{\text{NLO}} + imB_\mu\psi_{\text{LO}} - m_\mu\mathcal{D}_t\psi_{\text{LO}}$$

## QM on NR geometries

- Minimal coupling then leads to

$$\text{EOM}_{\text{LO}} = 2im\mathcal{D}_t\psi_{\text{LO}} + \mathcal{D}_i\mathcal{D}_i\psi_{\text{LO}} = 0$$

$$\text{EOM}_{\text{NLO}} = 2im\mathcal{D}_t\psi_{\text{NLO}} + \mathcal{D}_i\mathcal{D}_i\psi_{\text{NLO}} - \frac{1}{c^2}\mathcal{D}_t\mathcal{D}_t\psi_{\text{LO}} + \dots = 0$$

- Terms on the dots are fixed by demanding covariance under NLO diffeos ( $\zeta^i$ ) and residual diffeos ( $\xi^i$ ) of the LO geometry.
- The last step is to redefine  $\psi_{\text{NLO}}$  to a suitable  $\hat{\psi}_{\text{NLO}}$  in order that the Klein–Gordon inner product becomes the standard one:

$$\langle\varphi_{\text{KG}}|\psi_{\text{KG}}\rangle = \int_{t=\text{cst}} d^d x \left( \psi_{\text{LO}} + c^{-2}\hat{\psi}_{\text{NLO}} + \dots \right) \left( \varphi_{\text{LO}}^* + c^{-2}\hat{\varphi}_{\text{NLO}}^* + \dots \right)$$

## QM on NR geometries: Kerr geometry

$$ds_{\text{Kerr}}^2 = ds_{\text{flat}}^2 + \frac{\Sigma r_s R}{\Delta(R^2 + a^2)} dR^2 + \frac{r_s R}{\Sigma} (-cdt + a \sin^2 \Theta d\phi)^2$$

$$\Delta = R^2 + a^2 - r_s R, \quad \Sigma = R^2 + a^2 \cos^2 \Theta$$

- The parameters are  $r_s = \frac{2GM}{c^2}$  and  $a = \frac{J}{cM}$  with mass  $M$  and angular momentum  $J$  independent of  $c$ .
- $ds_{\text{flat}}^2$  is written in oblate spherical coordinates  $(R, \Theta, \phi)$ .
- Expanding in  $1/c^2$  and transforming to ordinary spherical coordinates  $(r, \theta, \phi)$  leads to the ‘Lense–Thirring metric’:

$$ds_{\text{Kerr}}^2 = - \left( 1 - \frac{2GM}{rc^2} + \frac{2GJ^2}{Mr^3c^4} P_2(\cos \theta) \right) c^2 dt^2 + \left( 1 + \frac{2GM}{rc^2} \right) dr^2 \\ + r^2 d\theta^2 + r^2 \sin^2 \theta d\phi^2 - \frac{4GJ}{rc^2} \sin^2 \theta dt d\phi + \mathcal{O}(c^{-4})$$

## QM on NR geometries: Kerr geometry

- Using the coupling prescription and defining the Hamiltonian  $H$  as  $i\partial_t\Psi = H\Psi$  where  $\Psi = \psi_{(0)} + c^{-2}\psi_{(2)} + \dots$  we find

$$H = \frac{p^2}{2m} - \frac{GmM}{r} - \frac{p^4}{8c^2m^3} + \frac{GM}{c^2m} \left( -\frac{3}{2r^3}x^i p_i x^j p_j + \frac{1}{2r^3}L^2 \right) - \frac{mG^2M^2}{2c^2r^2} + \frac{2GJ}{c^2r^3}L_z + \frac{mGJ^2P_2(\cos\theta)}{Mc^2r^3} + \frac{GM}{4mc^2}\Delta(r^{-1})$$

- This is the Hamiltonian of a spinless particle in a Kerr background up to order  $c^{-2}$ .
- This can be generalised to a spin 1/2 particle by starting with the Dirac equation.
- So far all top-down, but what if we want to study a hydrogen atom in a Kerr background?

Thank You!